

# The Many Mysteries of Cosmic Rays

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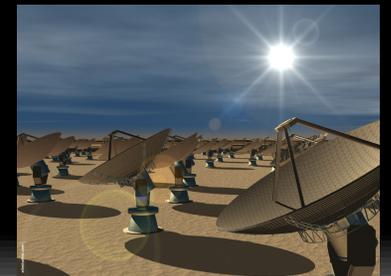
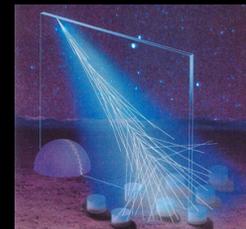
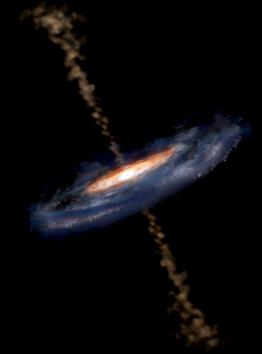


# Questions...

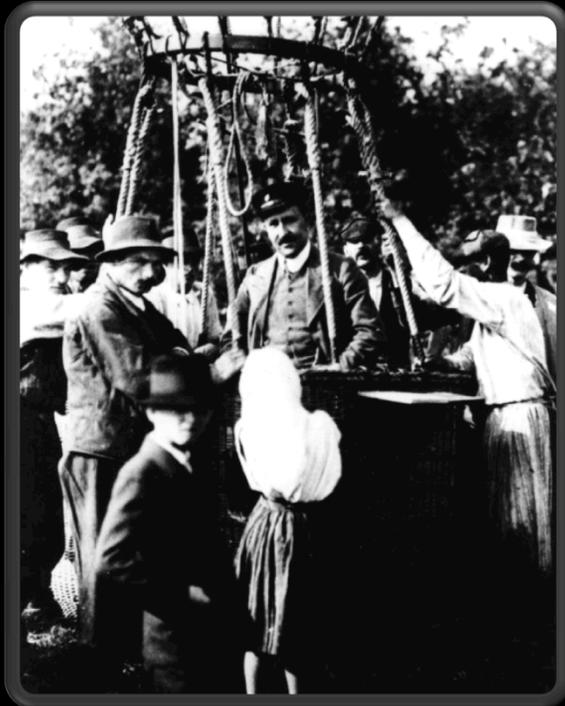
- What are cosmic rays?
- Where do they come from?

[ intermission? ]

- How can we detect cosmic rays?
- Where do we go from here?



# A bit of history



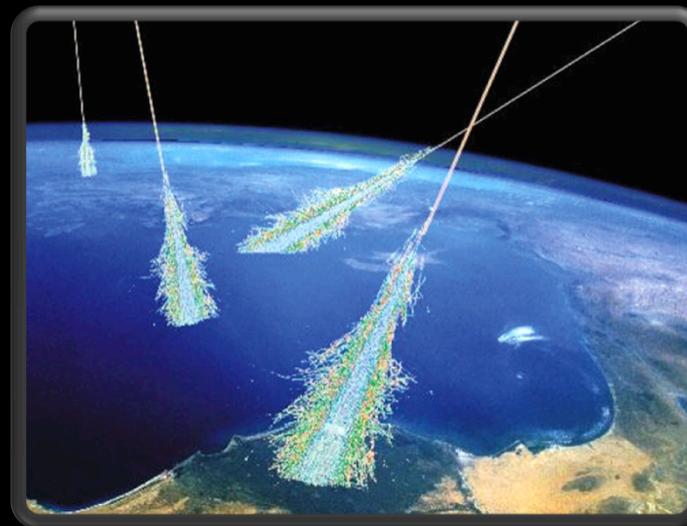
around 1900: scientists find mysterious background radiation (perhaps from the ground?)

1912: Victor Hess launches balloon — amount of radiation increases with height

*“... best explained by the assumption that a radiation of very great penetrating power enters our atmosphere from above”*

# What we know now...

- Cosmic “rays” are really charged particles from outer space
  - mostly protons, but also heavy nuclei...
- They hit the Earth from all directions in the sky, all the time
- Some are the highest energy particles known in the Universe



# What we *don't* know...

- Cosmic “rays” are really charged particles from outer space
  - mostly protons, but also heavy nuclei...

ALL of them? What about at the highest energies?

- They hit the Earth from all directions in the sky, all the time

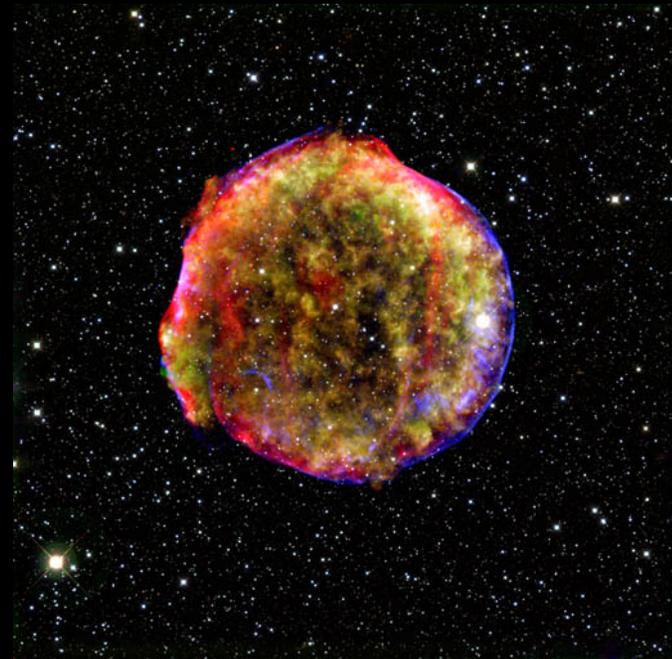
But where did they start?

- Some are the highest energy particles known in the Universe

How does Nature accelerate a particle to such high energies?

# Why study them?

- Might come from violent, interesting objects!
- Particle physics
  - 1933: discovery of antimatter
  - 1937: discovery of the muon
  - ...
  - 2008: hints of dark matter?
  - 2009+: testing Einstein's theory of relativity
- 2010+: Charged particle astronomy



Tycho's supernova remnant  
(from SN 1572, in Cassiopeia)

# Cosmic Ray Air Showers

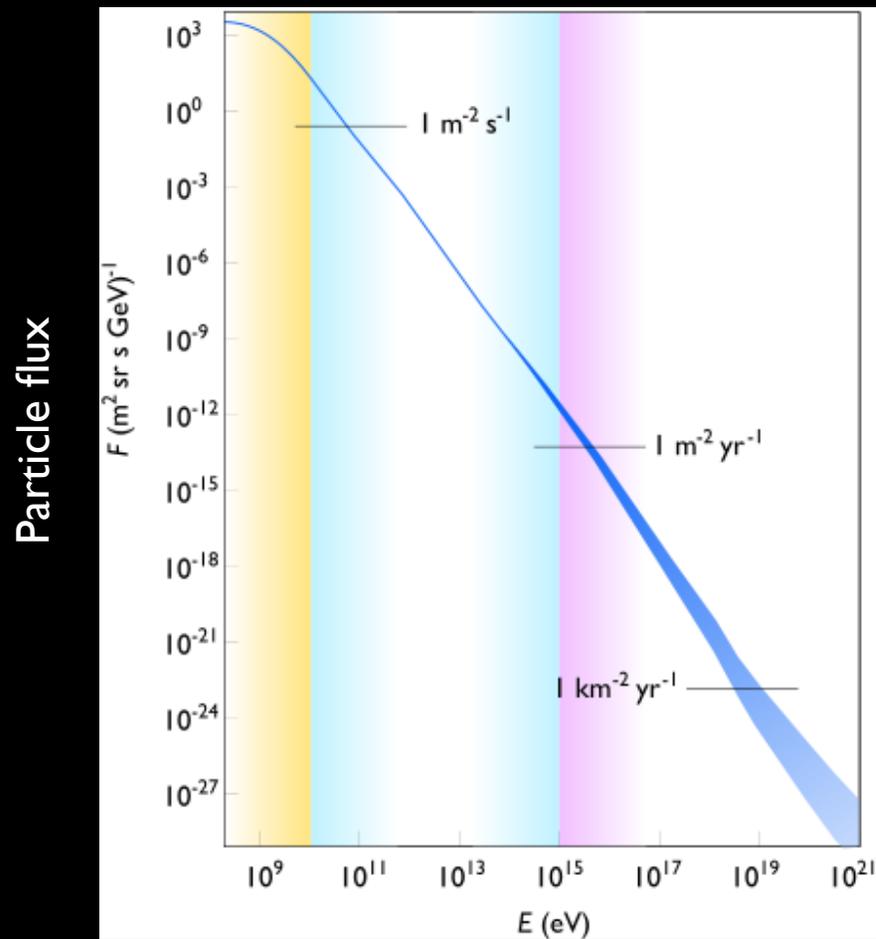
- Earth's atmosphere is our shield
- Chain reaction of particles
  - proton + air = pions (often)
  - “shower” of photons, electrons, positrons, muons, etc.
  - develops and hits ground in ~50 microseconds
- Detect particles that reach the ground
  - or look for flashes in the sky



# Air Shower Computer Simulation



# Energy Spectrum



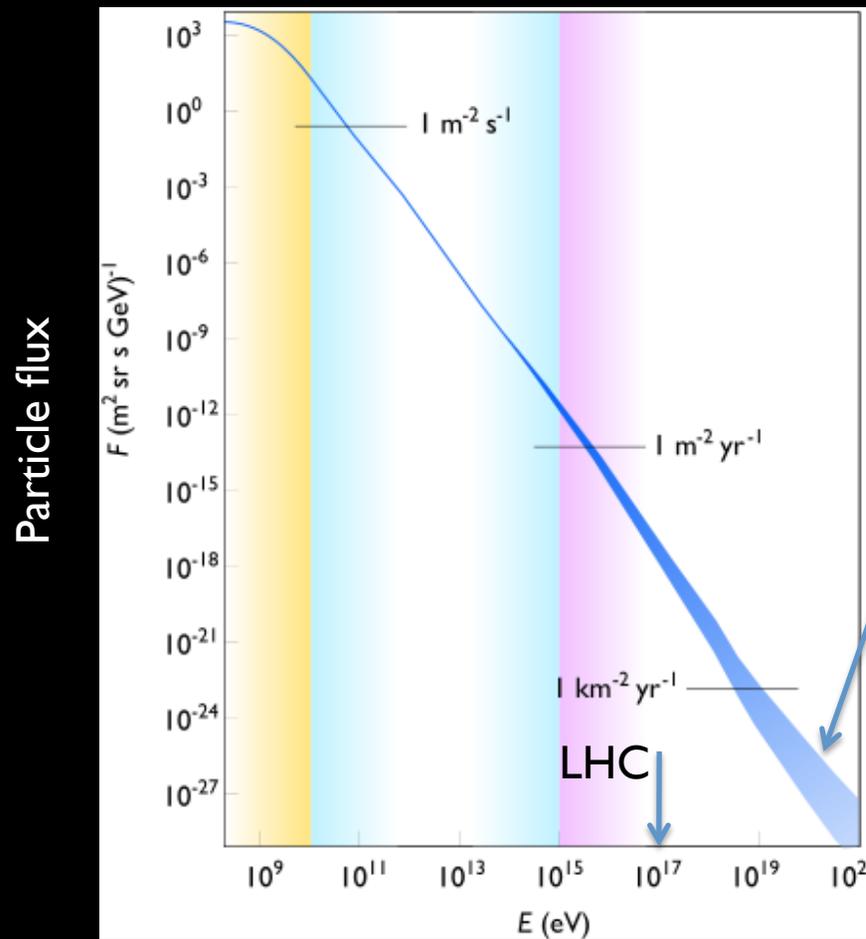
1 particle per  $\text{m}^2$   
per second

“Knee”:  
1 particle per  $\text{m}^2$  per year

“Ankle”:  
1 particle per  $\text{km}^2$  per year

Cosmic Ray Energy

# Energy Spectrum



Ultra-high energy  
cosmic rays:

1000x LHC energy  
but very rare  
(new particle physics?)

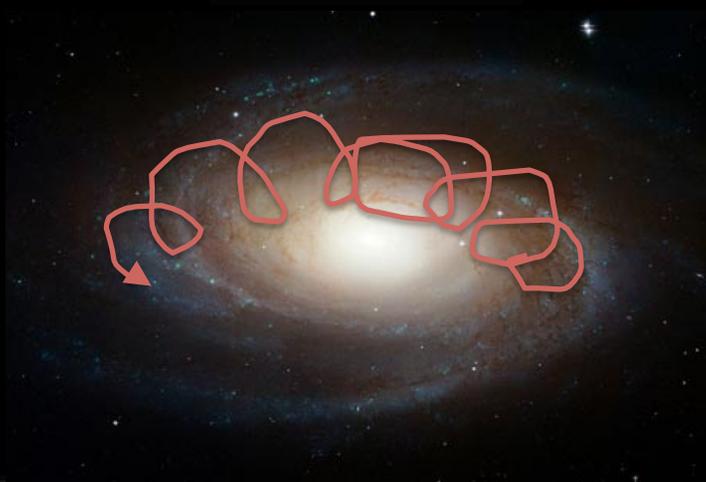
(P. S. why man-made  
particle colliders won't  
destroy Earth)

Cosmic Ray Energy

# Cosmic Ray Sources



- Main problem: magnetic fields scramble cosmic ray direction



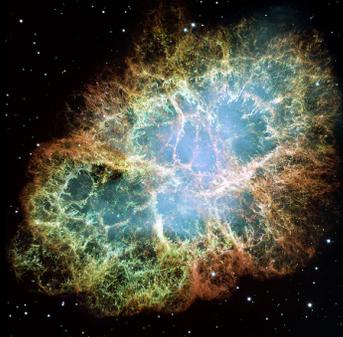
- May spiral around for a million years before hitting Earth

# One Galactic Possibility



- Enough supernovae (1 per ~50 years) to balance the energy budget
- Acceleration in shock wave (SNRs), not original explosion

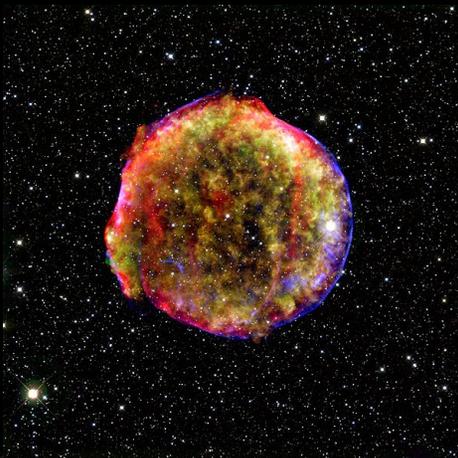
# Some SNRs



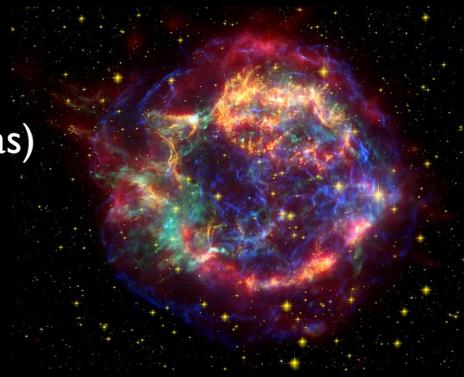
Crab nebula (M1)  
Hubble, false color



Veil nebula  
(NGC6992, etc.)  
optical

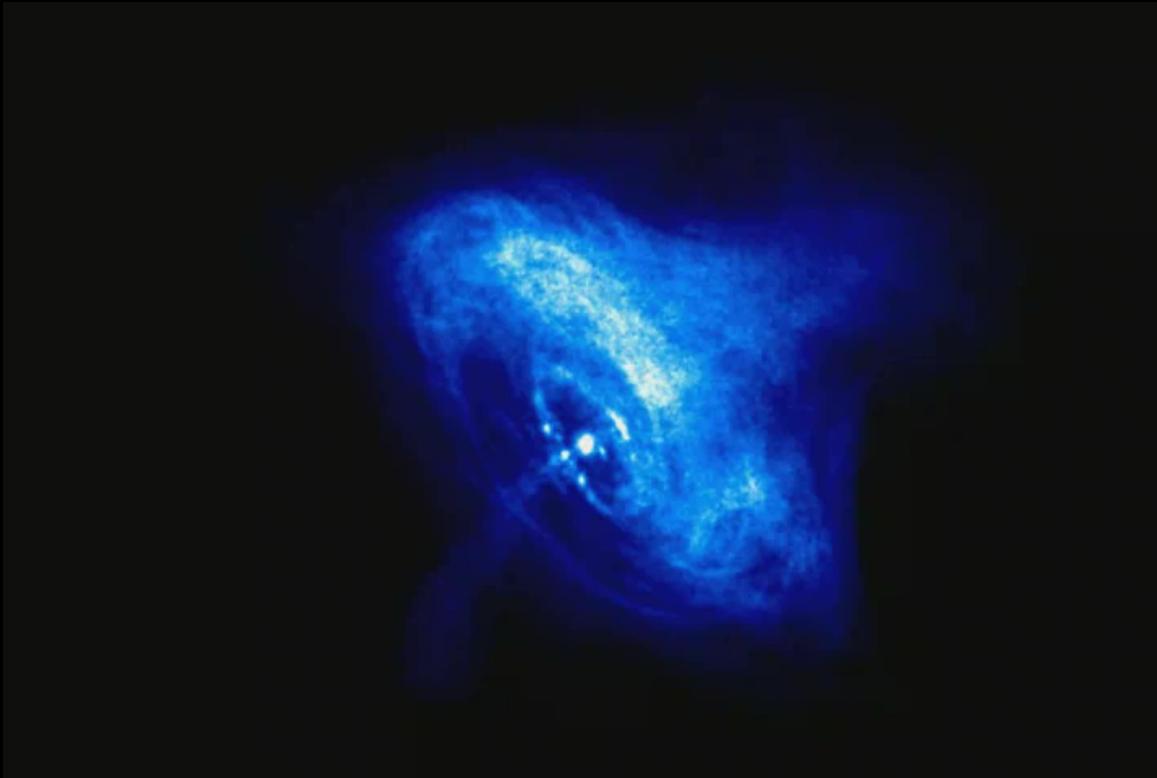


Tycho (SNI572 / B Cas)  
X-ray / infrared



Cassiopeia A  
X-ray / optical / infrared

# Experimental Status

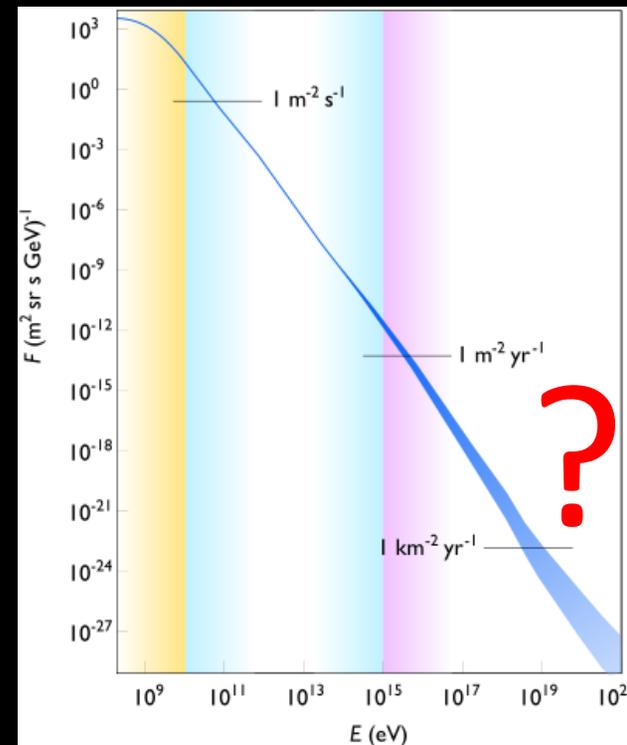


Chandra X-ray view of Crab core

- X-ray images show emission from high-energy particles in shocks
- Need independent confirmation by gamma ray and neutrino telescopes (no bending!)
- But only explains energies up to the “knee” (what about higher energies?)

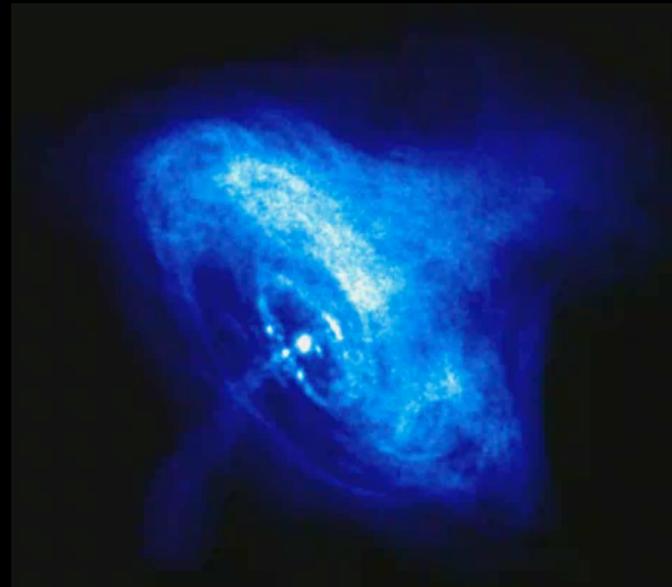
# Ultra-high Energy Cosmic Rays

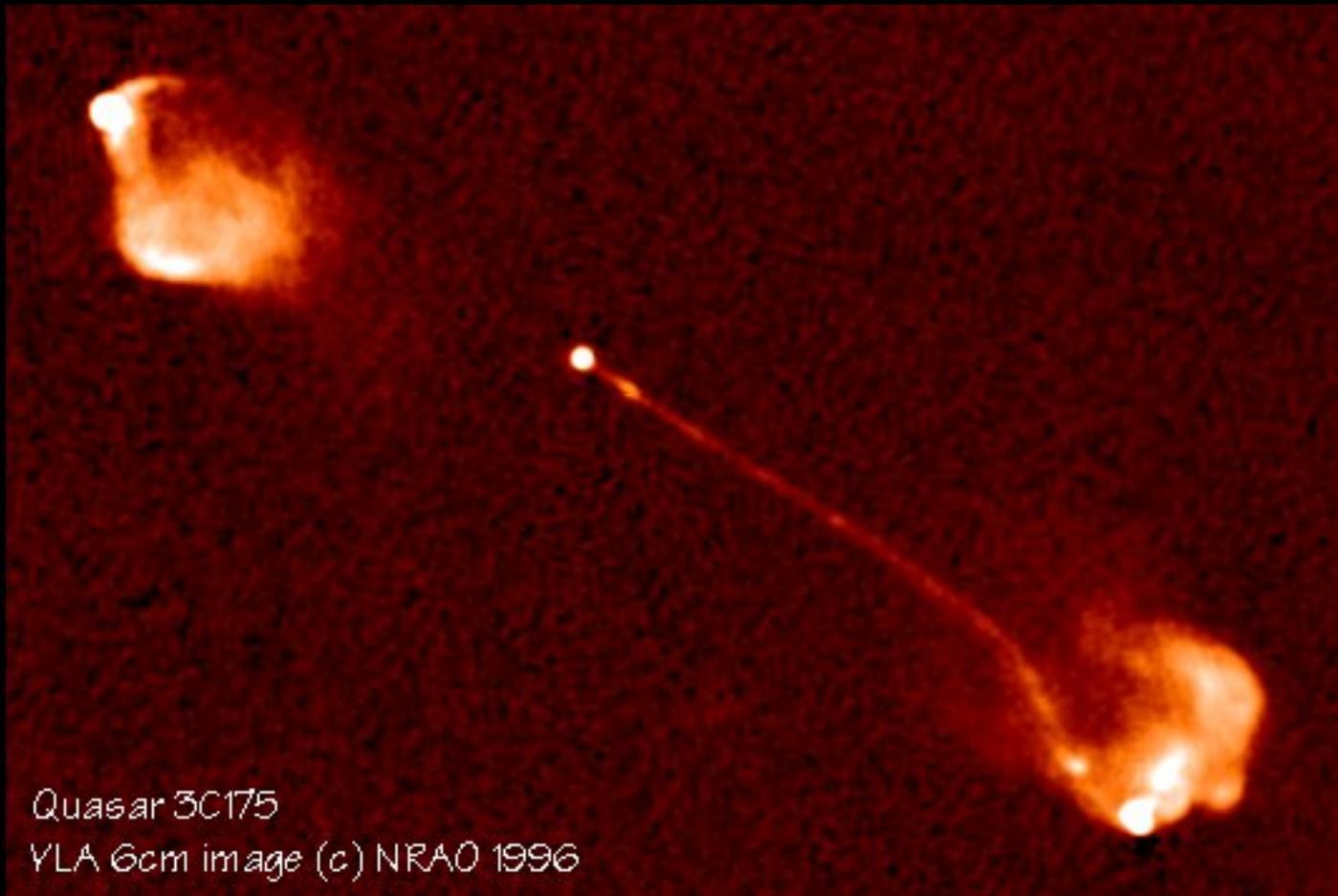
- Galactic magnetic fields too weak to hold them
- No known galactic sources can reach anywhere near  $10^{20}$  eV
- Probably come from outside our Galaxy



# Nature's Accelerators

- Rotating compact object (neutron star or black hole)
- Powers jets along spin axis via complex magnetic field
- Is there a REALLY BIG version?

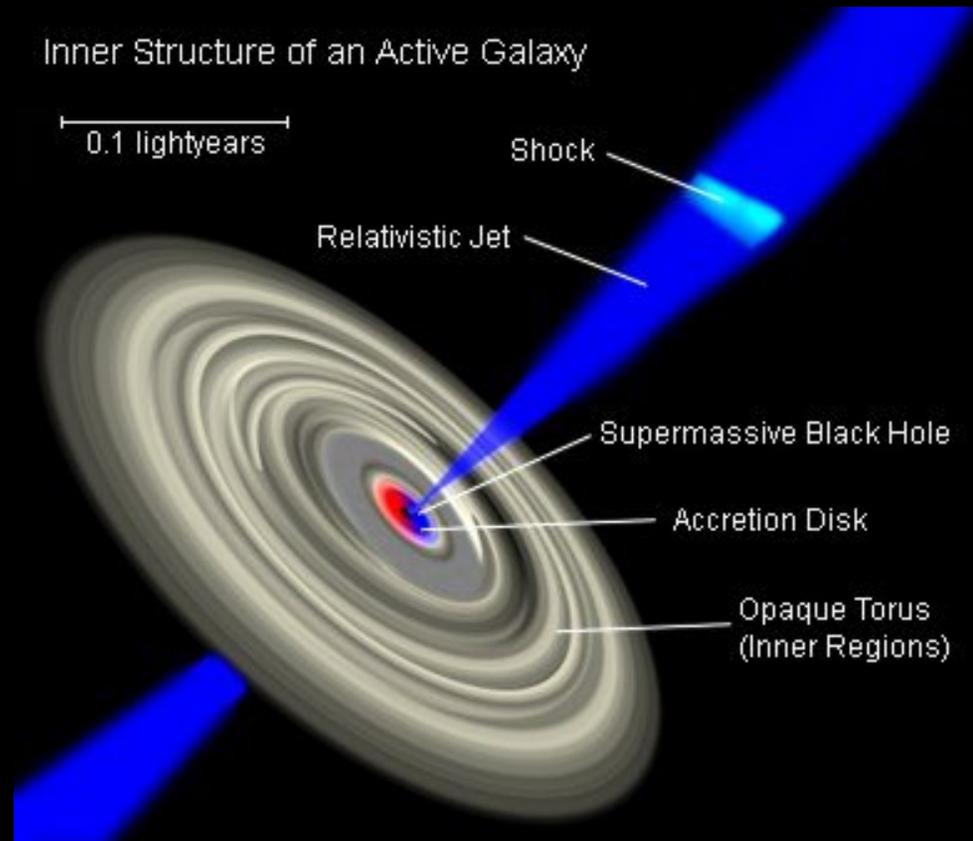




Quasar 3C175  
YLA 6cm image (c) NRAO 1996

# Active Galactic Nuclei (AGN)

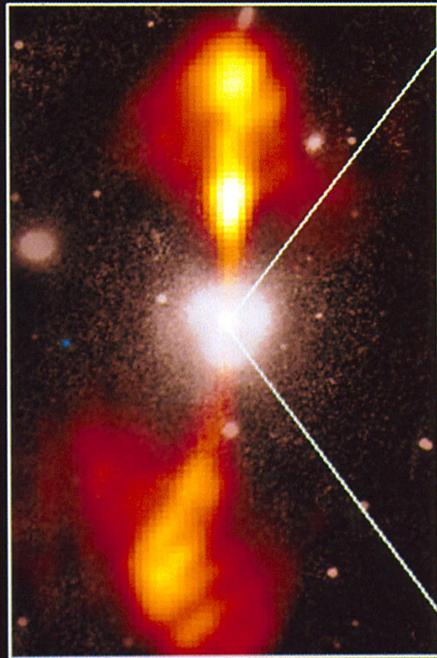
- Essentially all galaxies (including ours) house a supermassive black hole ( $10^6$  to  $10^{10}$  solar masses)
- Some are in an active phase (AGN)
- Shocks in jets: cosmic ray acceleration?



# Core of Galaxy NGC 4261

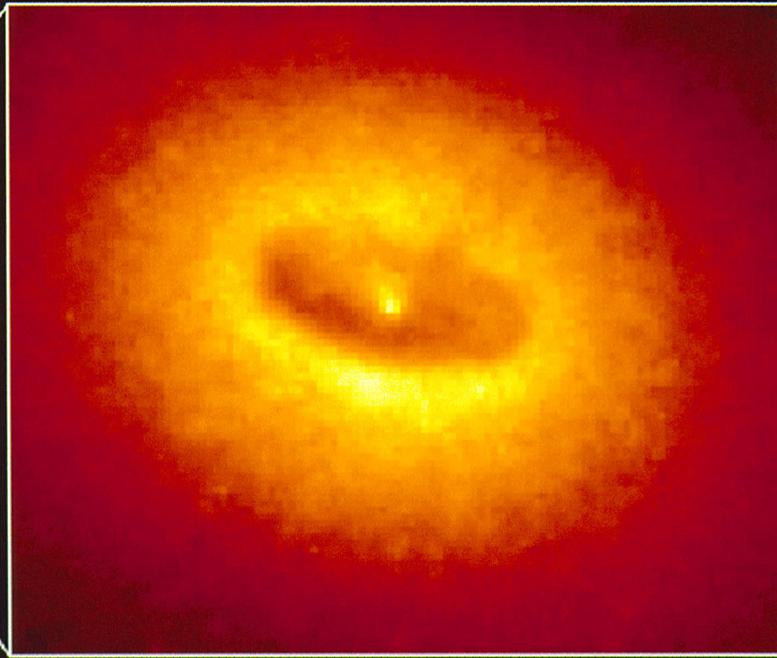
Hubble Space Telescope  
Wide Field / Planetary Camera

Ground-Based Optical/Radio Image



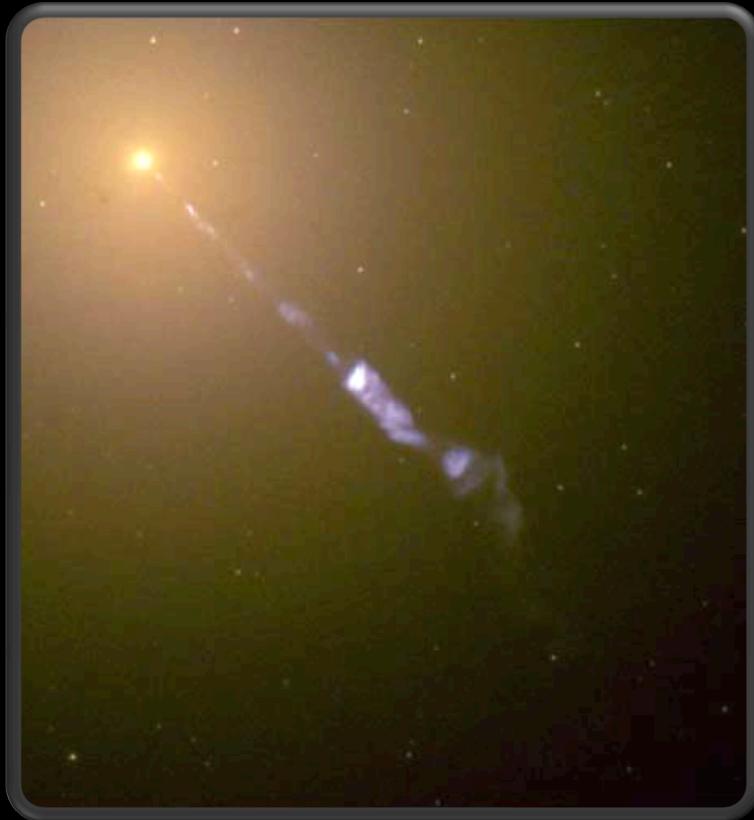
380 Arc Seconds  
88,000 LIGHT-YEARS

HST Image of a Gas and Dust Disk

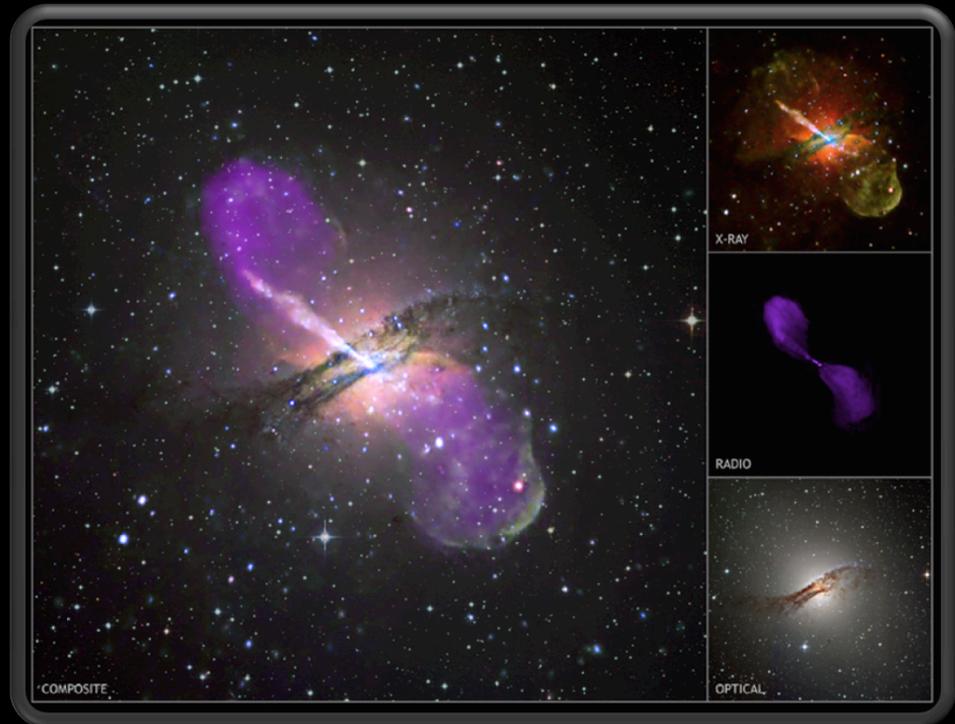


17 Arc Seconds  
400 LIGHT-YEARS

# M87

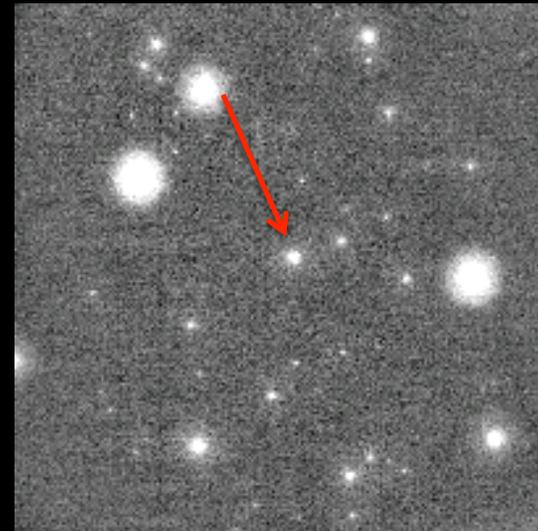


# Centaurus A (NGC 5128)



# March 19, 2008

- Light from an explosion nearly 8 billion years ago
- Visible with naked eye for 30 seconds in Boötes (most distant object ever)
- Spectacular example of a gamma ray burst (GRB 080319B)



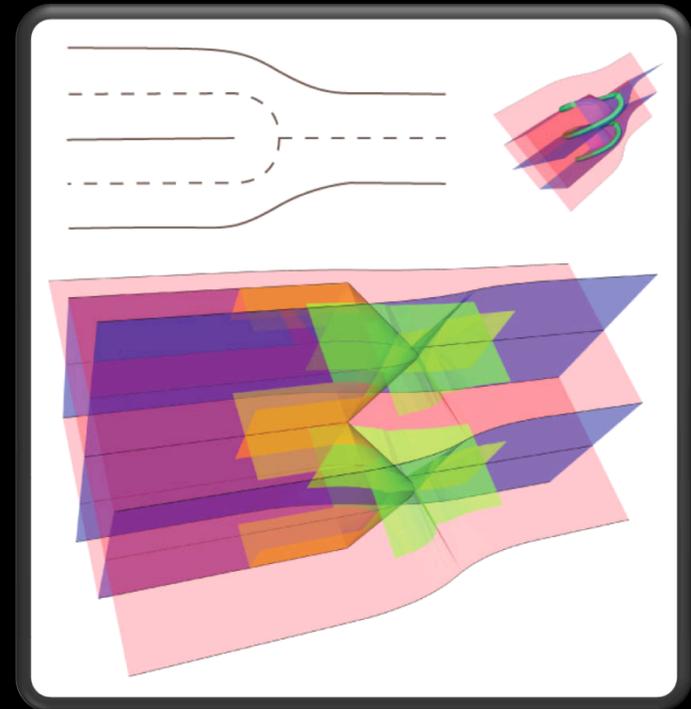
# GRB Animation



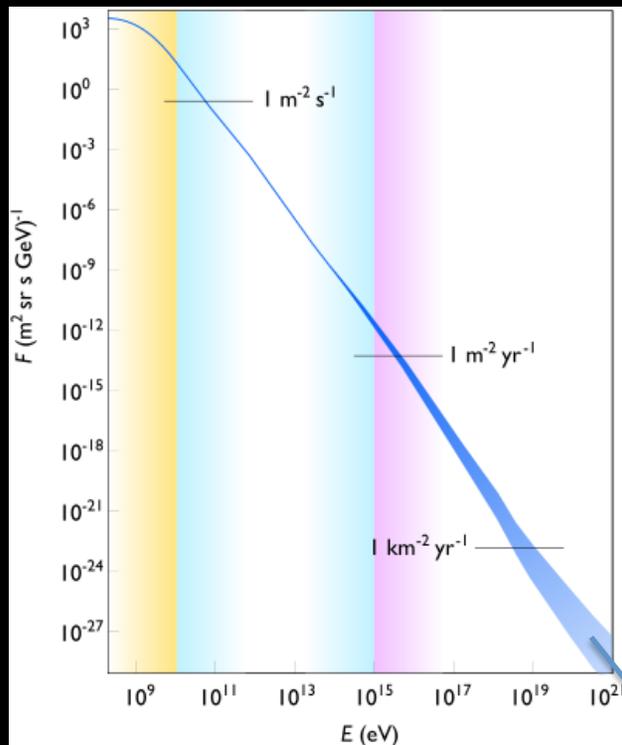
- Massive star (e.g. Wolf-Rayet) goes “hypernova”
- Inner core collapses to a black hole
- Jets bore through outer layers of star
- We see it if jet is pointed at us

# Other UHECR Sources

- AGN and GRBs: only known objects producing remotely enough energy
- Exotic ideas: decay of topological defects left over from Big Bang
- Something we haven't thought of?

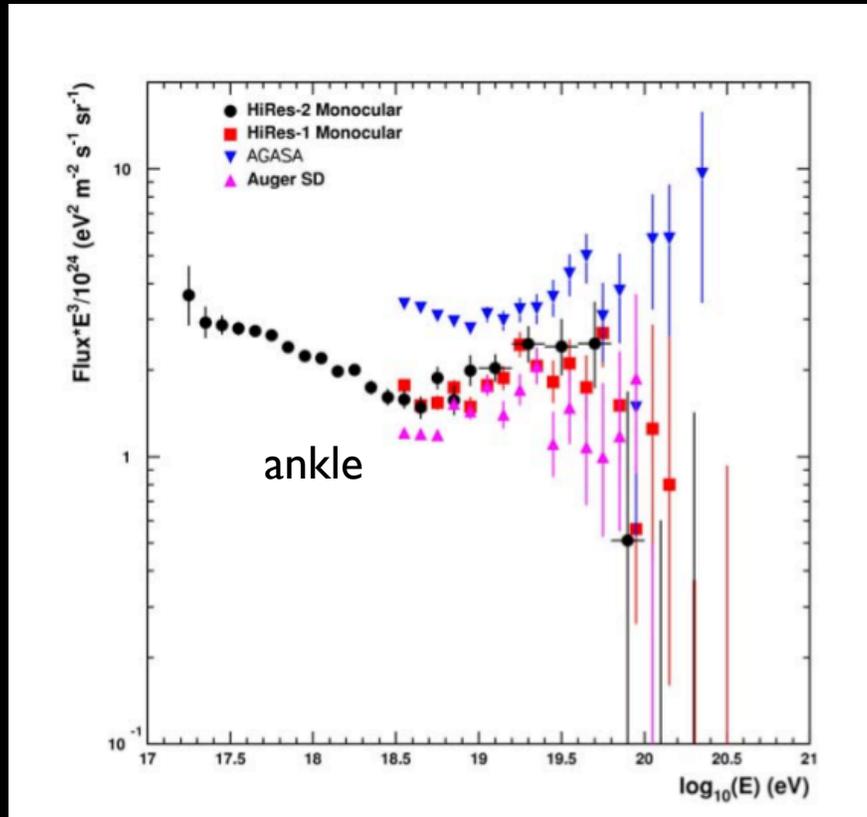


# GZK Cutoff



- Very high energy protons should react with glow from Big Bang (cosmic microwave background)  
 $p + \gamma \rightarrow \Delta^+$  (then decay)
- Expect spectrum to “cut off” above a certain energy (Greisen-Zatsepin-Kuzmin)
- If no cutoff:
  - sources are really close?
  - cosmic rays are not protons?
  - Einstein’s theory of special relativity is wrong (!)

# Situation in 2005



Cutoff? No cutoff?  
What's going on?!

Let's talk about detectors...

# A Bit of Particle Physics

One example

$$p + N \rightarrow \pi^+ + \text{fragments}$$

pion decay  $\pi^+ \rightarrow \mu^+ + \nu_\mu$

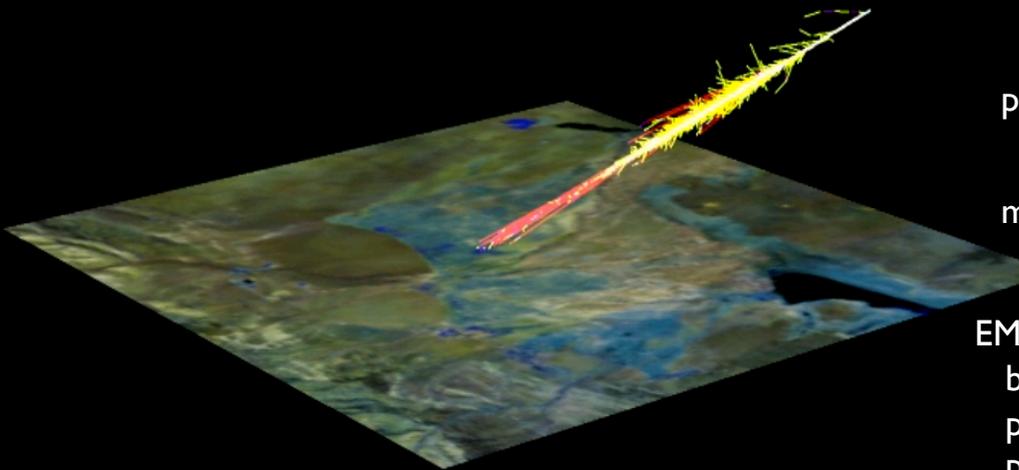
muon decay  $\mu^+ \rightarrow e^+ + \nu_e + \bar{\nu}_\mu$

EM shower:  
bremstrahlung  
pair creation  
pair annihilation

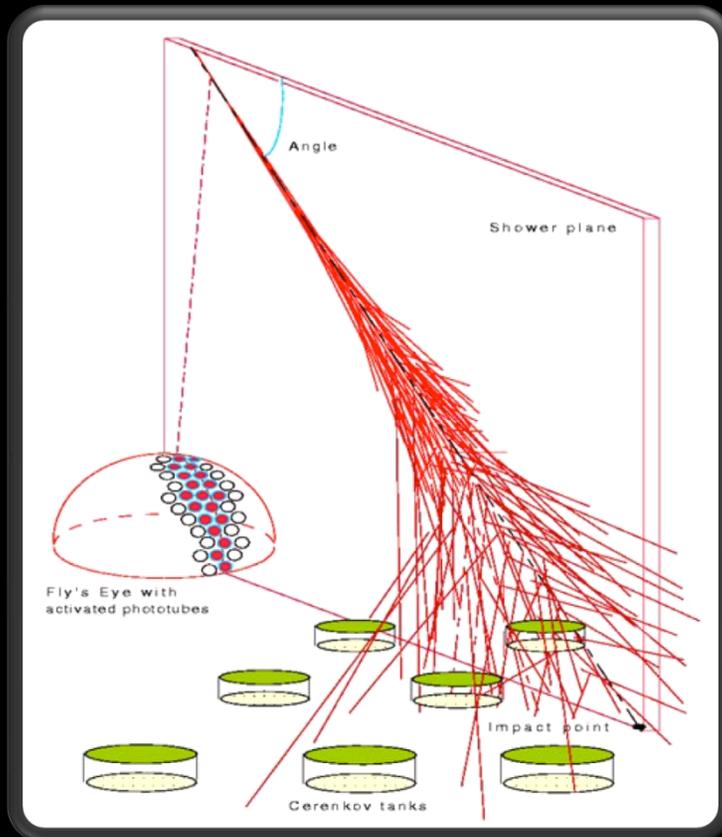
$$e^\pm + \gamma \rightarrow e^\pm + \gamma$$

$$\gamma + \gamma \rightarrow e^+ + e^-$$

$$e^+ + e^- \rightarrow \gamma + \gamma$$



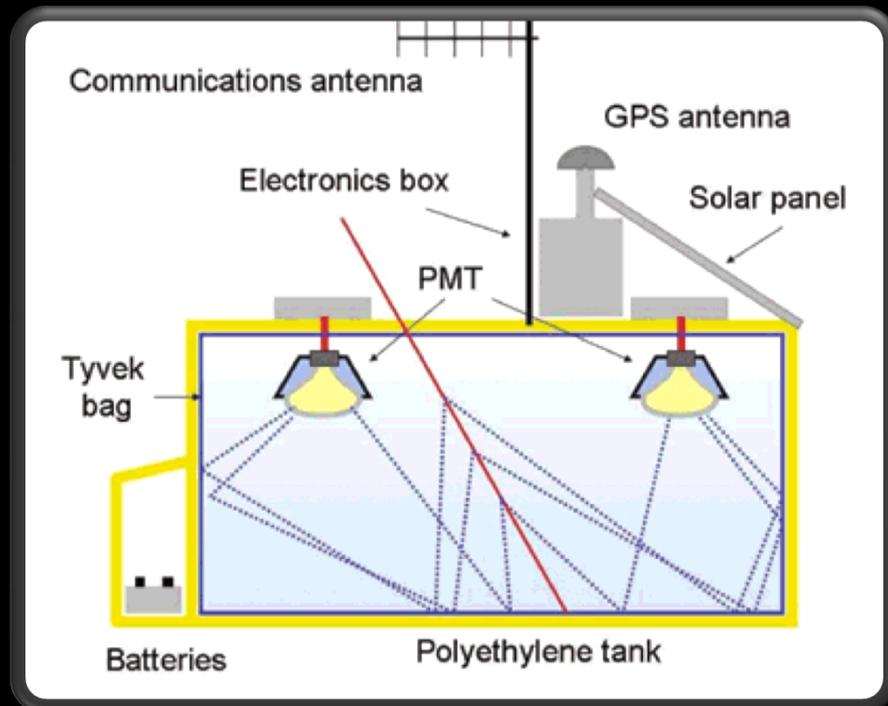
# Hybrid Air Shower Detector



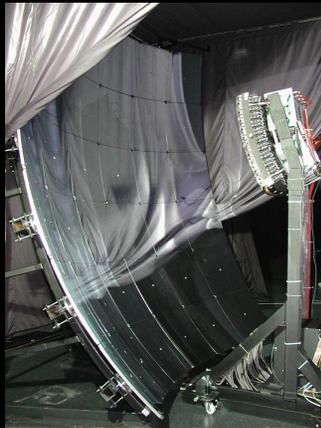
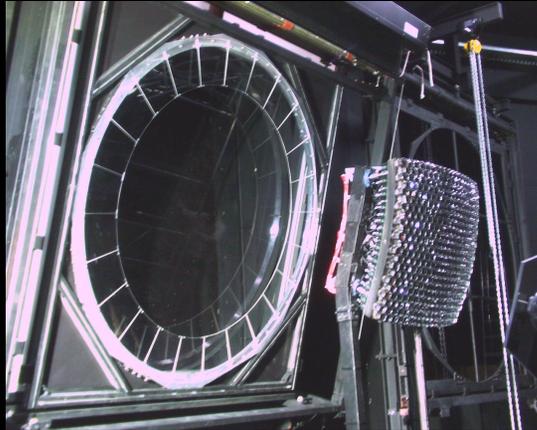
- Detect charged particles reaching ground with special water tanks
- See flash of light along track from fluorescence of air (if it's dark!)
- Each gives information about the cosmic ray (energy, particle type, etc.)
- Need huge area for highest-energy cosmic rays (or be really patient)

# Water Cherenkov Tanks

- Charged particle moving faster than speed of light in water (or any transparent medium): Cherenkov light
- Detect light inside tank with photomultiplier tubes (PMTs)
- Tank records signals along with an accurate time (from GPS)

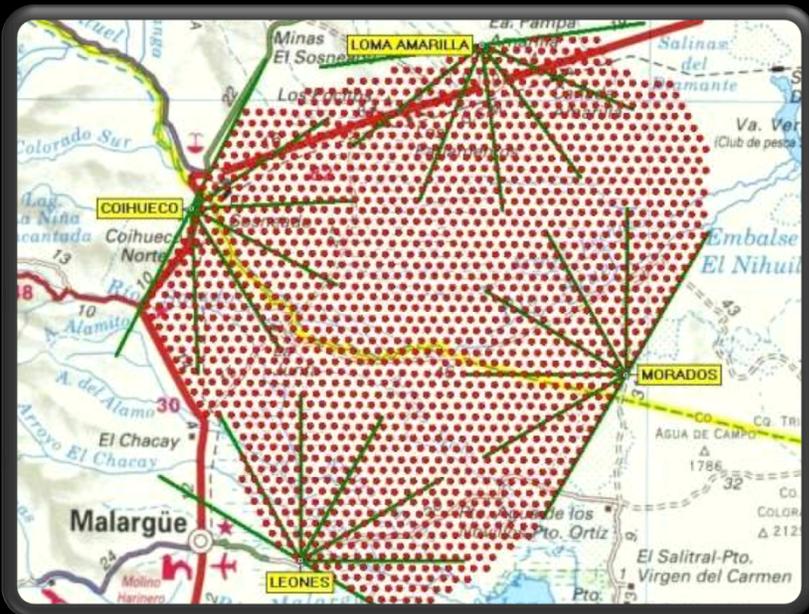


# Fluorescence Telescopes



- Air shower energy excites air molecules: fluorescence light
- Detect flash / track with array of small PMTs
- Must be dark, clear

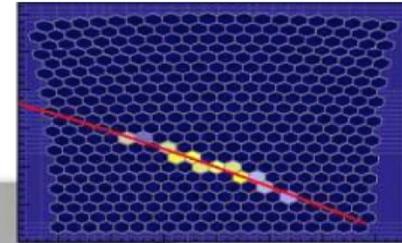
# The Pierre Auger Detector



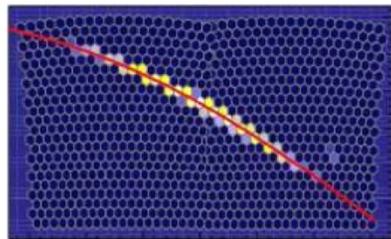
# Example Cosmic Ray Event

Event: 1364365

Los Morados

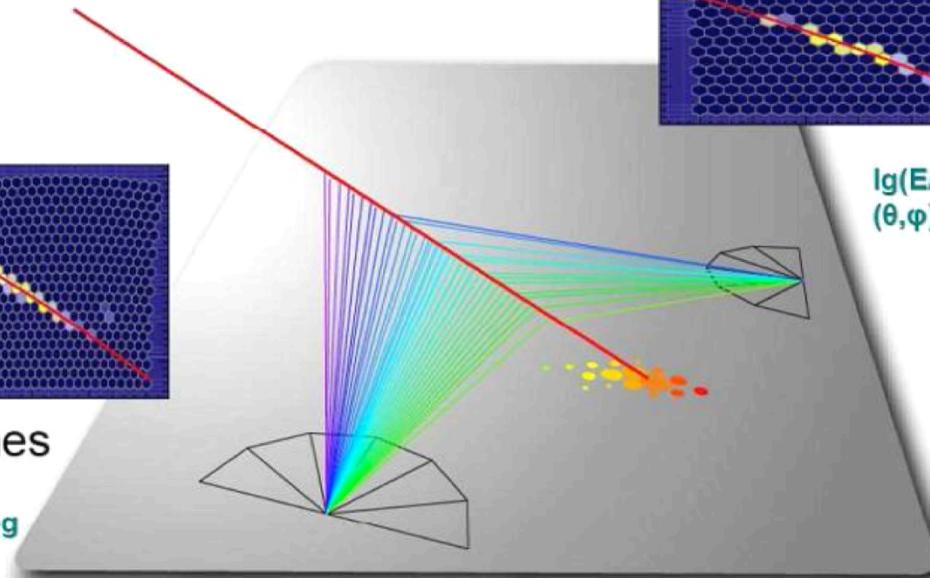


$\lg(E/eV) \sim 19.2$   
 $(\theta, \varphi) = (63.7, 148.4)$  deg



Los Leones

$\lg(E/eV) \sim 19.3$   
 $(\theta, \varphi) = (63.7, 148.3)$  deg



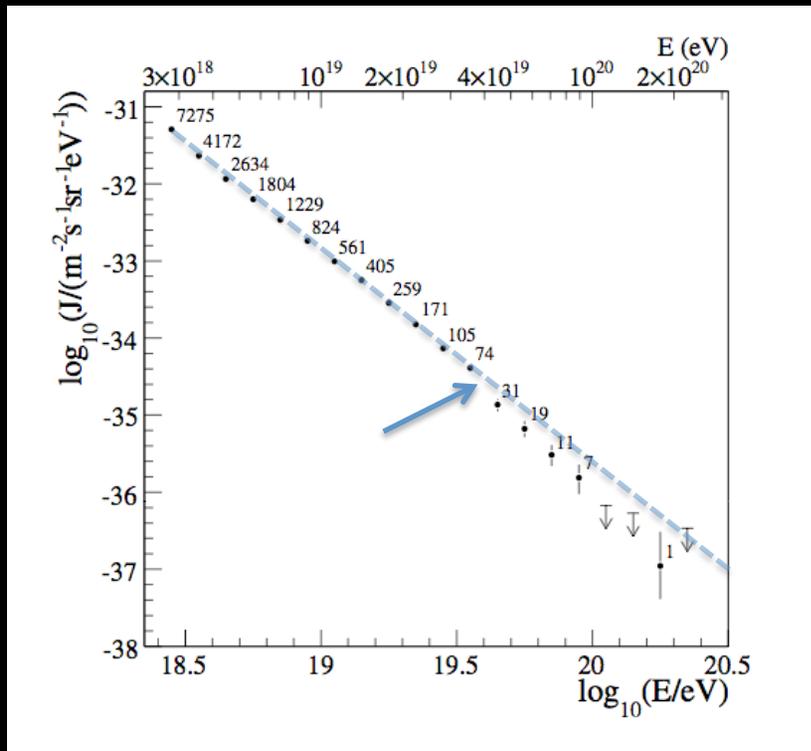
SD array:  $\lg(E/eV) \sim 19.1$   
 $(\theta, \varphi) = (63.3, 148.9)$  deg

$(\theta, \varphi) = (63.3, 148.9)$  deg

SD array:  $\lg(E/eV) \sim 19.1$

$(\theta, \varphi) = (63.3, 148.9)$  deg

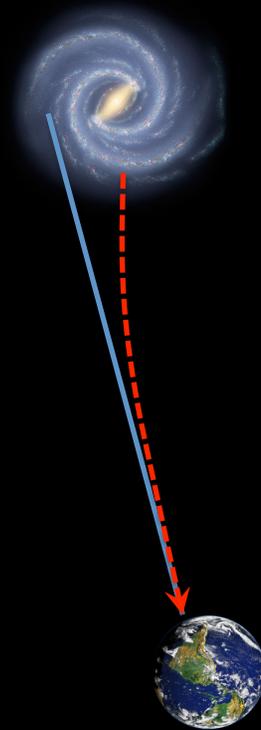
# Auger Energy Spectrum (2009)



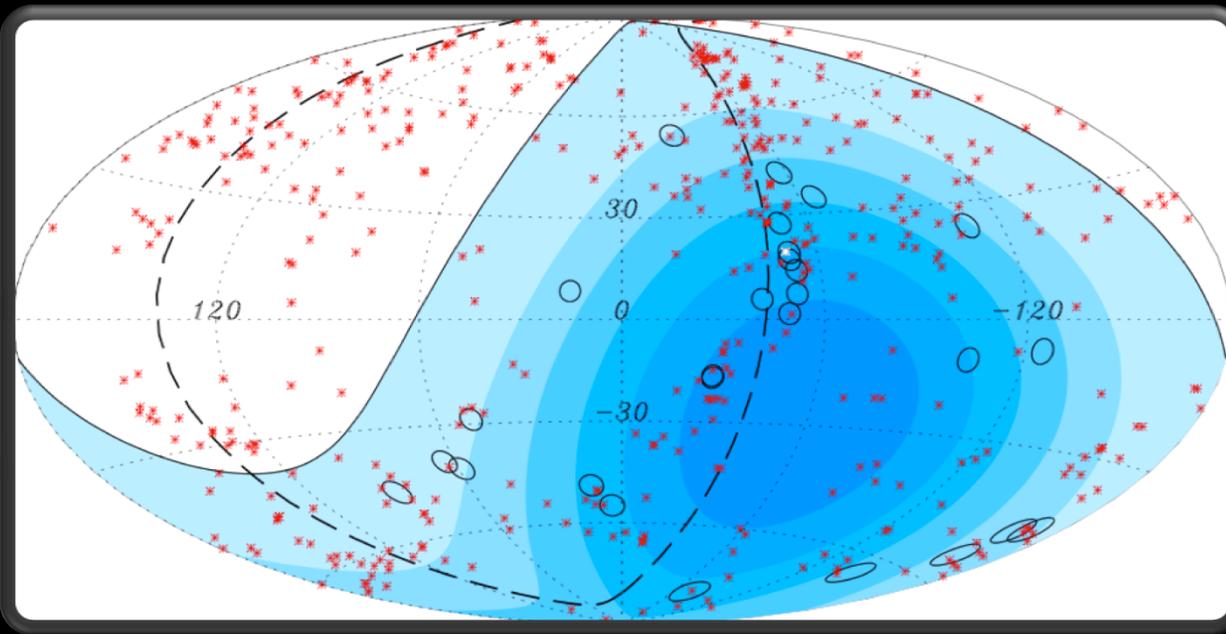
- “Break” in spectrum is where expected from GZK cutoff
- Statistics are not great (need more data)
- Future: shape of cutoff can tell us about distance of sources

# Auger UHECR Directions

- At very high energies, magnetic fields only bend protons by a few degrees
- Can point back to source?
- Cosmic rays closer to one type of object than some other type (correlation)?



# UHECR Skymap (2007)



- 38 highest-energy events (circles) seem to come from AGN (red dots)!
- Clump near Centaurus A (white dot)?
- BUT...
  - events since 2007 don't seem to correlate as well with AGN
- AND
  - events don't look like protons (curve more!)
- Need more data

# Satellites (or Balloons)



- In space, can detect the primary cosmic ray (instead of air shower)

BUT

- Spacecraft can only be around  $1 \text{ m}^2$

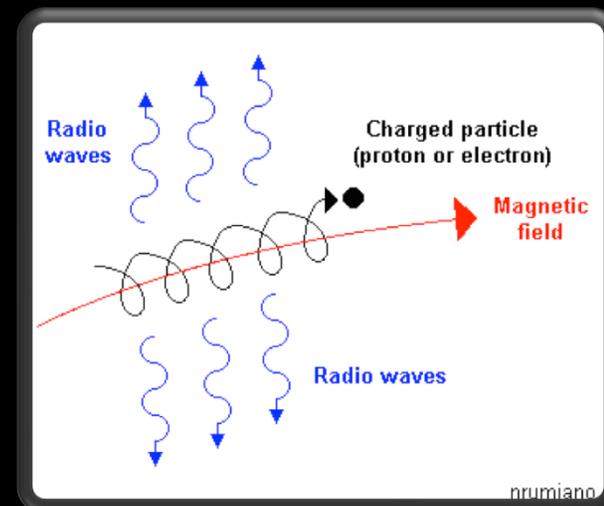
AND

- Balloon flights are short



# The Future: Radio Detection

- Electrons, positrons in air shower curve in Earth's magnetic field: emit radio pulse (geosynchrotron)
- Idea: detect cosmic ray air showers with radio antennas
- Complements other detection methods

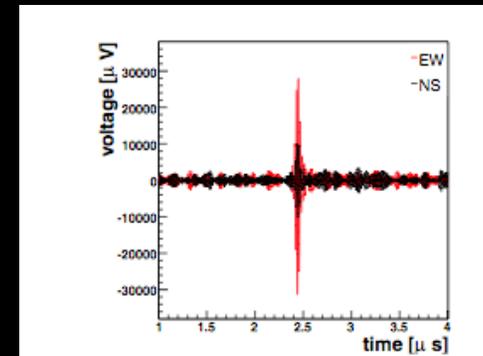


# Prototype Station

antenna

electronics

solar panels



An air shower pulse!

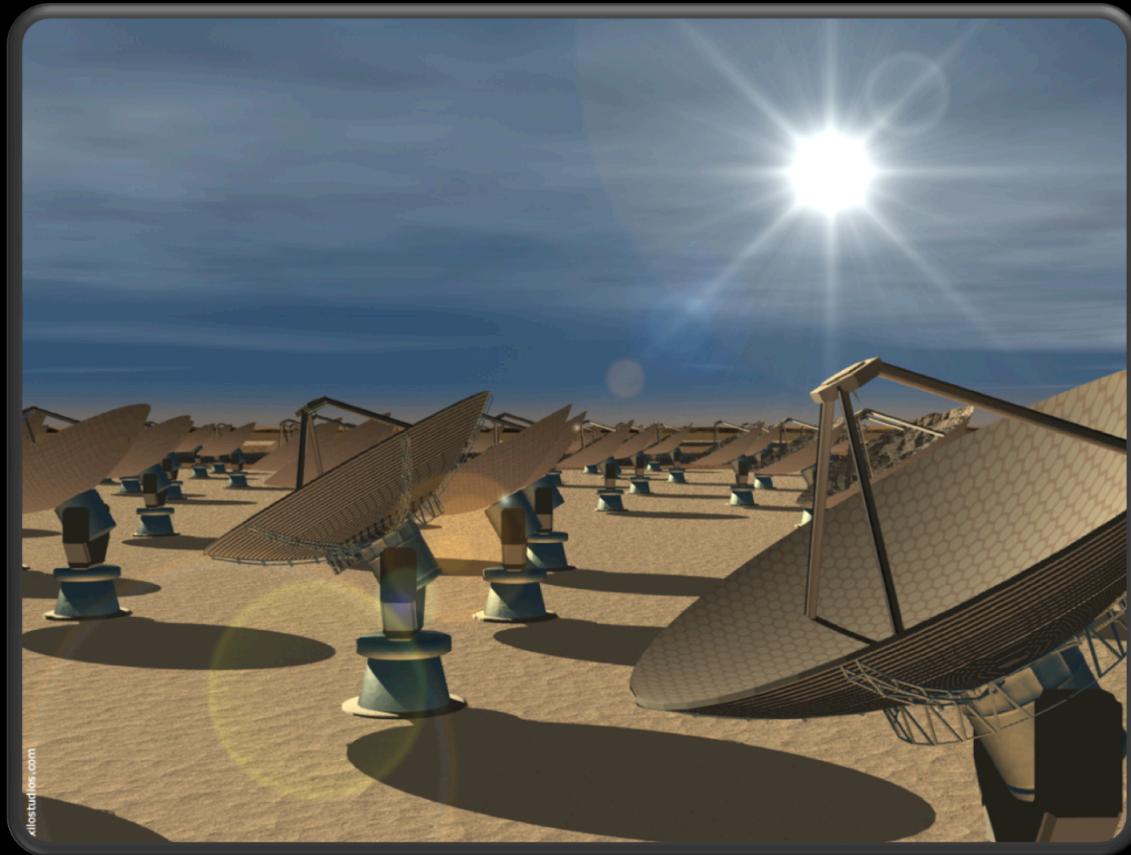
(also radio stations, TV stations,  
cars, power lines...)



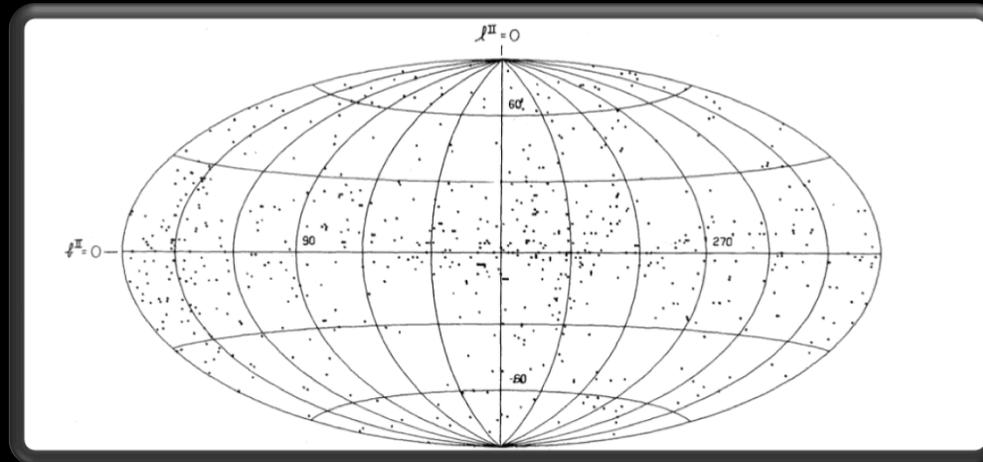
# Next-generation Radio Telescopes



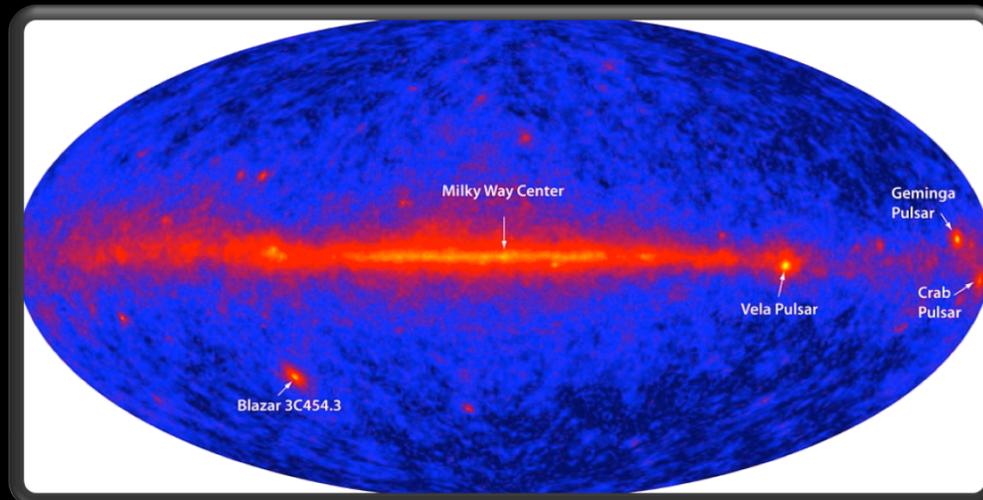
# Square Kilometer Array



# An Analogy with Gamma Ray Astronomy

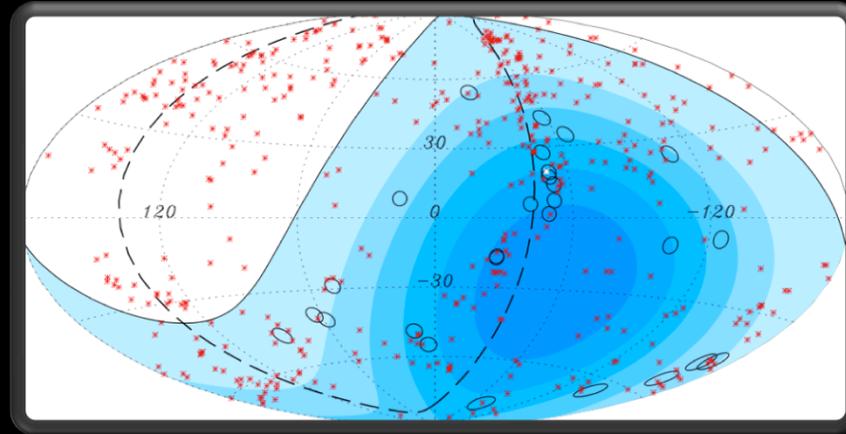


OSO-3 (1968)



Fermi (2008)

# Charged Particle Astronomy



Auger (2007)

???

The future?

# Summary

- Still many unanswered questions about cosmic rays
  - May come from SNRs and AGN... not sure
  - Many are protons, but at high energies...?
- Large air shower detectors like Auger continue taking data
  - We should know a lot more in a few years
- New experimental techniques like radio air shower detection in development
- The beginning of a new type of astronomy!



SN 1006 remnant (in Lupus)

Thank you!